3/11/13 Lecture outline

• Recall  $A^{\mu} = (\phi, \vec{A}), j^{\mu} = (c\rho, \vec{j}).$  Lorentz force law:

$$
\frac{dp^{\mu}}{d\tau} = f^{\mu} = \gamma \frac{dp^{\mu}}{dt} = \gamma (q\vec{E} + \frac{q}{c}\vec{v} \times \vec{B}) = \frac{q}{c}F^{\mu\nu}u_{\nu}.
$$

The time component gives the power:  $\gamma \frac{d\mathcal{E}}{dt}$ . Maxwell's equations:  $\partial_{\mu}F^{\mu\nu} = \frac{4\pi}{c}$  $\frac{1}{c}J^{\nu}$  and  $\partial_{\mu}\widetilde{F}^{\mu\nu}$ , where  $\widetilde{F}^{\mu\nu}=\frac{1}{2}$  $\frac{1}{2} \epsilon^{\mu\nu\rho\sigma} F_{\rho\sigma}$ , or equivalently  $\partial_{\mu} F_{\rho\sigma} + \partial_{\rho} F_{\sigma\mu} + \partial_{\sigma} F_{\mu\rho} = 0$ ; we solve the latter via  $F_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu}$ . Using transformation  $F_{\mu'\nu'} = \Lambda_{\mu}^{\mu'}\Lambda_{\nu}^{\nu'}F^{\mu\nu}$ , find

$$
\vec{E}_{||} = \vec{E}_{||}^{\prime}, \qquad \vec{E}_{\perp} = \gamma (\vec{E}^{\prime} - \frac{\vec{v}}{c} \times \vec{B}^{\prime})_{\perp}
$$

$$
\vec{B}_{||} = \vec{B}_{||}^{\prime}, \qquad \vec{B}_{\perp} = \gamma (\vec{B}^{\prime} + \frac{\vec{v}}{c} \times \vec{E}^{\prime})_{\perp}
$$

• Moving point charge:  $J^{\mu} = c\rho \frac{dx^{\mu}}{dx^0}$ , which is a 4-vector because  $\rho$  and  $dx^0$  transform the same way. Likewise,  $\rho = q\delta^3(\vec{x} - \vec{x}_0)$  makes sense, with q Lorentz invariant. E.g.  $\delta^4(x^{\mu}-x_0^{\mu})$  is Lorentz invariant, and  $\delta(t-t_0)dt$  is Lorentz invariant. The term  $-\frac{q}{c}A_{\mu}dx^{\mu}$ in the point particle world-line action can thus be written as a spacetime volume integral − 1  $\frac{1}{c} \int d^4x A_\mu J^\mu$ , which is Lorentz invariant. Note  $d^4x$  is Lorentz invariant and  $\epsilon^{\mu\nu\rho\sigma}$  is Lorentz invariant for the same reason, mentioned last time:  $\det \Lambda = 1$ .

• Lorentz invariants:  $F_{\mu\nu}F^{\mu\nu} = 2(B^2 - E^2)$  and  $F_{\mu\nu}\tilde{F}^{\mu\nu} = -4\vec{E}\cdot\vec{B}$ . Note that if  $\vec{E} \cdot \vec{B} = 0$ , then there is a frame where the field is entirely  $\vec{E}'$  or entirely  $\vec{B}'$ , depending on the sign of  $F_{\mu\nu}F^{\mu\nu}$ .

• Maxwell's equations as a field theory:

$$
S_{field} = \int d^4x \mathcal{L}_{field}, \qquad \mathcal{L}_{field} = -\frac{1}{16\pi} F_{\mu\nu} F^{\mu\nu}
$$

where we impose  $F_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu}$ . The field interacts with charges via

$$
S_{int} = -\frac{1}{c} \int d^4x A_\mu J^\mu.
$$

Varying  $A^{\mu} \rightarrow A^{\mu} + \delta A^{\mu}$  and requiring that the action be stationary gives Maxwell's equations.

• The total action is  $S = S_{field} + S_{matter} + S_{int}$ . Spacetime translation symmetry,  $x^{\mu} \to x^{\mu} + \epsilon^{\mu}$  is related to conservation of  $P^{\mu} = (H, c\vec{P})$ . We saw that electric charge conservation is equivalent to  $\partial_{\mu}J^{\mu}=0$ . Likewise, conservation of  $P^{\mu}$  is

$$
P^{\mu} = \int d^3x T^{\mu 0} \quad \text{conserved} \quad \leftrightarrow \quad \partial_{\nu} T^{\mu \nu} = 0.
$$

The relation between the conservation law and the symmetry is Noether's theorem:

$$
\frac{d}{dx_{\mu}}\mathcal{L} = \frac{d}{dx^{\nu}}\left(\frac{\partial \mathcal{L}}{\partial(\partial_{\nu}A_{\lambda})}\partial^{\mu}A_{\lambda}\right) + \frac{\partial \mathcal{L}}{\partial x_{\mu}}
$$

which implies

$$
\partial_{\nu}T^{\mu\nu} = -\frac{\partial \mathcal{L}}{\partial x_{\mu}}, \qquad T^{\mu\nu}_{field} = \frac{\partial \mathcal{L}_{field}}{\partial(\partial_{\nu} A_{\lambda})} \partial^{\mu} A_{\lambda} - g^{\mu\nu} \mathcal{L}_{field}.
$$

So if there is no explicit  $x_{\mu}$  dependence then the conservation equation for  $T^{\mu\nu}$  is satisfied. We'll show next time how this gives the field contribution to energy and momentum densities, and how  $T_{tot}^{\mu\nu} = T_{matter}^{\mu\nu} + T_{field}^{\mu\nu}$ , with (only) the sum conserved.