

$\mathcal{L} = \int d^4\theta K(\Phi_i, \bar{\Phi}_{\bar{i}})$  has susy vacua with

○ arbitrary  $\langle \phi_i \rangle \in \Sigma$  with metric

$$g_{\Sigma} = \partial_i \bar{\partial}_{\bar{j}} K(\phi_i, \bar{\phi}_{\bar{i}}). \text{ "Non linear sigma model"}$$

Global  $U(1)_R$  symmetry ( $\sim$  fermion number)

$\phi_i, \bar{\phi}_{\bar{i}} =$  neutral,  $\psi_i$  charge  $-1$ ,  $\psi_{\bar{i}}^+$  charge  $+1$ .

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To lift  $\langle \phi_i \rangle$  degeneracy, add a scalar potential  $V(\phi_i, \bar{\phi}_{\bar{i}})$ , e.g.  $\phi_i$  mass terms. In susy

○ theory this comes from adding a Superpotential:

$$\mathcal{L} = \int d^4\theta K(\Phi_i, \bar{\Phi}_{\bar{i}}) + \int d^2\theta W(\Phi_i) + \int d^2\bar{\theta} \bar{W}(\bar{\Phi}_{\bar{i}})$$

susy requires  $W$  to be holomorphic:  $\frac{\partial W}{\partial \bar{\Phi}_{\bar{i}}} = 0$

and  $\bar{W}$  to be anti-holomorphic  $\frac{\partial \bar{W}}{\partial \Phi_i} = 0$

in the chiral superfields.

$$\text{Reality: } \bar{W}(\bar{\Phi}_{\bar{i}}) = [W(\Phi)]^+$$

$$\int d^2\theta W \rightarrow \partial_i W F^i - \frac{1}{2} \partial_i \partial_{\bar{j}} W \psi^i \psi^{\bar{j}}$$

$$F^i \text{ EOM: } F^i = \frac{1}{2} \Gamma_{jk}^i \psi^j \psi^k - g^{i\bar{j}} \partial_{\bar{j}} \bar{W}$$

get

$$\mathcal{L} = g_{i\bar{j}} \left( -\partial_{\bar{r}} \phi^i \partial^{\bar{r}} \bar{\phi}^{\bar{j}} + \bar{\psi}^{\bar{j}} (iD) \psi^i \right)$$

$$+ \frac{1}{24} R_{i\bar{j}k\bar{l}} \psi^i \psi^k \bar{\psi}^{\bar{j}} \bar{\psi}^{\bar{l}}$$

$$- \frac{1}{2} (\nabla_i \nabla_{\bar{j}} W) \psi^i \psi^{\bar{j}} + \text{h.c.} \leftarrow \text{"Yukawa terms"}$$

$$- V(\phi, \bar{\phi})$$

$V \geq 0$  as expected by susy

with scalar potential  $V(\phi, \bar{\phi}) = g^{i\bar{j}} \partial_i W \partial_{\bar{j}} \bar{W}$ .

Supersymmetric vacuum requires  $\langle F^i \rangle = 0$

$$\frac{1}{\epsilon} \langle \psi^i \rangle = 0 \quad \frac{1}{\epsilon} \dots \langle g^{i\bar{j}} \partial_{\bar{j}} \bar{W} \rangle = 0$$

this gives  $\langle V(\phi, \bar{\phi}) \rangle = 0$  in a susy vacuum

as expected since  $\langle H \rangle = 0$  if

$Q_\alpha$  &  $\bar{Q}_{\dot{\alpha}}$  annihilate the vacuum.

For non degenerate  $g_{ij}$  the condition for a  
 ○ Susy vacuum is  $\therefore$  that there is a  
 solution  $\langle \phi_i \rangle$  to the eqns  $\frac{\partial W}{\partial \phi_i} = 0$

note this is  $n$  eqns in  $n$  unknowns  $\langle \phi_i \rangle$   
 for  $i = 1 \dots n$ .

If  $\langle F^i \rangle \neq 0$ , susy is broken &  
 $\psi^i \sim$  massless Goldstino

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○ Examples take one field  $\Phi$  &  $K = \bar{\Phi} \Phi$

1)  $W = \lambda \Phi \rightarrow F = -\lambda^* \neq 0$  susy broken  
 no sol'n to  $\frac{dW}{d\Phi} = 0$  for  $\lambda \neq 0$ .

$$\mathcal{L} = -\partial_\mu \bar{\Phi} \partial^\mu \Phi - i \bar{\Psi}_\alpha \partial^{\alpha\alpha} \Psi_\alpha - |\lambda|^2$$

free theory with energy offset:  $V = |\lambda|^2 \neq 0$   
 like a cosmological constant. Breaks susy.

○  $\Psi_\alpha \rightarrow$  massless Goldstino  $\text{Tr}(-1)F = 0$ .

2)  $W = \frac{m}{2} \overline{\Phi}^2 \rightarrow$  the following terms in  $\mathcal{L}$ :

$$-W'' \psi \psi = -m \psi \psi \leftarrow \psi \text{ gets mass } m \quad \circ$$

$$-|W'|^2 = -m^2 |\phi|^2 \leftarrow \text{complex scalar } \phi \text{ gets mass } m$$

Same mass, as required by susy.

Susy vacuum requires  $W' = 0$  i.e.  $\langle \phi \rangle = 0$

This vacuum has  $\text{Tr}(-1)^F = 1$ .

Recall in susy QM with 2 supercharges

$$\text{Tr}(-1)^F = \text{sign}(m) \text{ with } m = \text{real} \quad \circ$$

$$\text{Here 4 supercharges } \hat{Q} \quad \text{Tr}(-1)^F = \text{sign}(mm^*) = 1$$

Since  $m = \text{complex}$ , can rotate  $m \rightarrow e^{i\alpha} m \hat{Q}$

$\text{Tr}(-1)^F$  must be unchanged under continuous deformations

So  $\text{Tr}(-1)^F = 1$  for any  $m \neq 0$ .

$$3) W = \frac{m}{2} \overline{\Phi}^2 + \lambda \overline{\Phi} = \frac{m}{2} \frac{\tilde{\Phi}}{m}^2 - \frac{\lambda^2}{2m}$$

$$\frac{\tilde{\Phi}}{m} \equiv \overline{\Phi} + \lambda/m$$

Susy vacuum at

$$\langle \tilde{\Phi} \rangle = 0$$

$\nearrow$   
constant term in  $W$   
has no effect in global susy.

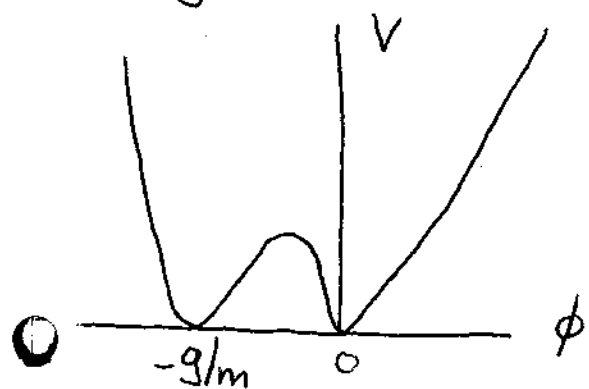
(Does contribute in SUGRA)

$$4) W = \frac{g}{3} \Phi^3 + \frac{1}{2} m \Phi^2 \rightarrow \text{fermion } \mathcal{L} =$$

$$\circ -W'' \psi \psi = -(2g\phi + m) \psi \psi \leftarrow \begin{array}{l} \text{mass + Yukawa} \\ \text{interaction} \end{array}$$

$$-|W'|^2 = -|m\phi + g\phi^2|^2 = -V(\phi, \bar{\Phi})$$

$$2 \text{ susy vacua : } \phi = 0 \quad \& \quad \phi = -g/m$$



Each vacuum contributes  $(-1)^F = 1$ , total  $\text{Tr}(-1)^F = 2$ .

$$5) W = g \Phi^{n+1} \rightarrow \text{Susy vacua at}$$

$$W'(\phi) = 0 \rightarrow \langle \phi \rangle = 0. \text{ Actually } n \text{ vacua}$$

on top of each other. Defining  $W$

$$\text{e.g. } W = g \Phi^{n+1} + \lambda \Phi \text{ or general}$$

order  $n+1$  polynomial.  $W' =$  order  $n$  polynomial

$\circ$  in complex  $\phi \rightarrow \exists$   $n$  solutions to

$$W'(\langle \phi \rangle) = 0. \text{ Each has } (-1)^F = 1 \text{ so total}$$

$$\text{Tr}(-1)^F = n. \quad n \text{ susy vacua.}$$

$$6) K = (\bar{\Phi} \Phi)^{1/2} \quad W = \lambda \bar{\Phi}$$

can do change of variables:  $X \equiv \bar{\Phi}^{1/2}$   $\circ$

to get canonical  $K = \bar{X} X \stackrel{!}{=} W = \lambda X^2$

so must have  $\text{Tr} (-1)^F = 1$  susy vacuum at  $\langle X \rangle = 0$ . How to see:  $\bar{\Phi} = \dots + \theta^2 F$

$$\text{so } \int d^4\theta K \rightarrow \dots + \frac{\partial^2 K}{\partial \phi \partial \bar{\phi}} F \bar{F} = \frac{(\phi \bar{\phi})^{-1/2}}{4} F \bar{F}$$

$\int d^2\theta W \rightarrow \lambda F$ . So  $F$  EOM  $\rightarrow$   $\circ$

$$\bar{F} = 4 \lambda (\phi \bar{\phi})^{1/2} \stackrel{!}{=} V(\phi, \bar{\phi}) = |F|^2 = 16 |\lambda|^2 |\phi|^2$$

can satisfy  $\langle F \rangle = 0 \stackrel{!}{=} \langle V \rangle = 0$  at  $\phi = 0$   $\checkmark$

place where  $\frac{\partial^2 K}{\partial \phi \partial \bar{\phi}}$  coeff of kinetic terms

is singular. Singularity  $\Rightarrow$  should change

variables to nonsingular field  $X$  to  $\circ$

see correct physics.

More comments on  $U(1)_R$  symmetries. Assign  $\Theta^\alpha$

charge +1  $\frac{1}{2}$ ;  $\bar{\Theta}^{\dot{\alpha}}$  charge -1

○

$$\mathbb{I}(y, \Theta) = \phi(y) + \sqrt{2} \Theta \psi(y) + \Theta^2 F(y)$$

If  $\mathbb{I}$  has charge  $q$ ,  $\phi$  charge =  $q$ ,

$\psi$  charge =  $q-1$ ,  $F$  charge =  $q-2$ .

$$\mathcal{L} = \int d^4\theta K + \int d^2\theta W + \text{h.c.}$$

respects  $U(1)_R$  provided  $K$  charge = 0  $\frac{1}{2}$

○  $U(1)_R$  charge of  $W = 2$ .

E.g.  $K = \bar{\Phi} \Phi$   $\frac{1}{2}$ ;  $W = g \bar{\Phi}^{n+1}$  respects

$U(1)_R$  with  $\bar{\Phi}$  charge =  $2/n+1$ . Adding

other terms to  $W$  would break this  $U(1)_R$

e.g.  $W = \frac{g}{3} \bar{\Phi}^3 + \frac{m}{2} \bar{\Phi}^2$  has no

unbroken  $U(1)_R$  if both  $g \neq 0$   $\frac{1}{2}$ ;  $m \neq 0$ .

○ We can pretend that this breaking is

"spontaneous" rather than explicit by

Thinking of  $g$  &  $m$  as constant expectation values of chiral superfields rather than as parameters. E.g.  $W = \frac{g}{3} \overline{\Phi}^3 + \frac{m}{2} \overline{\Phi}^2$

$U(1)_R$  charge

$$\overline{\Phi} \quad 2/3$$

$$g \quad 0$$

$$m \quad 2/3$$

← "field"  $m$  is charged under  $U(1)_R$ . Its vev  $\langle m \rangle \neq 0$

∴ Spontaneously breaks  $U(1)_R$ .

This viewpoint will be very important shortly.

Another comment about  $U(1)_R$ : R charge of  $W = 2$ .

Scaling dimension of  $W = 3$ . In a scale inv.

Susy theory (= superconformal field theory)

there is a conserved  $U(1)_R$  such that

all chiral superfields  $\overline{\Phi}$  have

$$\text{Scaling dim} [\overline{\Phi}] = \frac{3}{2} \text{ R charge} [\overline{\Phi}].$$



Another example: two chiral superfields  $\underline{\Phi}_1, \underline{\Phi}_2$

○ with  $K = \overline{\Phi}_1 \Phi_1 + \overline{\Phi}_2 \Phi_2$  &  $W = \frac{1}{2} g \overline{\Phi}_1 \Phi_2^2$ .

$$V(\phi_1, \overline{\phi}_1) = \left| \frac{\partial W}{\partial \phi_1} \right|^2 + \left| \frac{\partial W}{\partial \phi_2} \right|^2 = \left| \frac{1}{2} g \phi_2^2 \right|^2 + |g \phi_1 \phi_2|^2$$

Susy vacua when  $\frac{\partial W}{\partial \phi_1} = 0$  &  $\frac{\partial W}{\partial \phi_2} = 0$

$$\rightarrow \langle \phi_2 \rangle = 0 \quad \langle \phi_1 \rangle = \text{arbitrary.}$$

○ There is a (classical) "moduli" space of supersymmetric vacua for any  $\langle \phi_1 \rangle \in \mathbb{C}$ .

For  $\langle \phi_1 \rangle \neq 0$ , the superpotential for  $\underline{\Phi}_2$  is  $W \approx \frac{1}{2} g \langle \phi_1 \rangle \overline{\Phi}_2^2 \rightarrow$  mass for  $\underline{\Phi}_2$

$$= g \langle \phi_1 \rangle. \quad \text{Since } \underline{\Phi}_2 \text{ is massive,}$$

we can "integrate it out" for energies

○ below its mass. Do path integral  $\int [d\underline{\Phi}_2]$

to get effective field theory for  $\underline{\Phi}_1$ .

Low energy quantum effective theory for  $\underline{\Phi}_1$ ,

$$\mathcal{L}_{\text{eff}} = \int d^4\theta K_{\text{eff}}(\underline{\Phi}_1, \overline{\underline{\Phi}}_1) + \int d^2\theta W_{\text{eff}}(\underline{\Phi}_1) + \text{h.c.}$$

Since  $\langle \underline{\Phi}_2 \rangle = 0$ ,  $W_{\text{cl}}(\underline{\Phi}_1) = 0$

so there is the classical moduli space of vacua mentioned earlier. Is this classical degeneracy lifted by quantum effects? I.e. is

a  $W_{\text{eff}}(\underline{\Phi}_1) \neq 0$  generated?

No! Original theory has  $U(1)_R$  symmetry

with  $\underline{\Phi}_1 = \text{neutral}$   $\hat{=}$   $\underline{\Phi}_2$  charge = 1

$\langle \underline{\Phi}_1 \rangle \neq 0$  leaves this symmetry unbroken.

So any  $W_{\text{eff}}(g, \underline{\Phi}_1)$  must respect this  $U(1)_R$  i.e. have charge 2. But can not make charge 2  $W_{\text{eff}}$  from neutral  $g \hat{=}$   $\underline{\Phi}_1$ .  $\therefore W_{\text{exact}}(g, \underline{\Phi}_1) = 0$

There is an exact quantum moduli space of degenerate  $\text{susy vacua}$   $\langle \underline{\Phi}_1 \rangle = \text{arb.}$

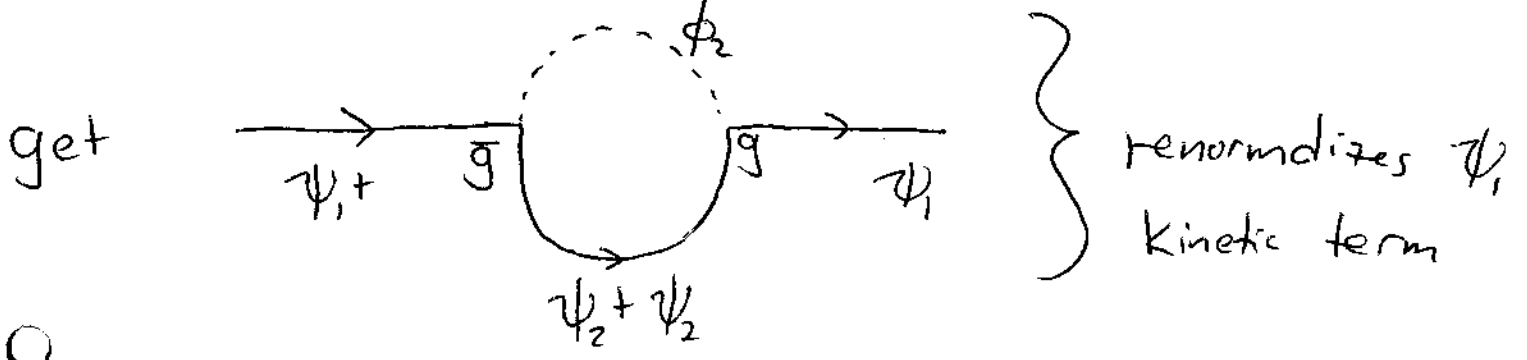
$$R_{\text{eff}} = \int d^4\theta K_{\text{eff}}(\overline{\Phi}_1, \Phi_1) + \int d^2\theta W_{\text{eff}}(\Phi_1) + \text{h.c.}$$

○ nontrivial quantum effects in h.c. ↑

○ ↓

e.g.  $W_{\text{eff}}(\overline{\Phi}_1, \Phi_2) \quad k = \frac{1}{2} g \overline{\Phi}_1 \Phi_2^2$  leads to

Yukawa interactions  $-g \phi_2 \psi_1 \psi_2 - g \phi_1 \psi_2 \psi_2 + \text{h.c.}$



○  $\psi_1^+ (i\partial) \psi_1 \rightarrow Z^2 \psi_1^+ (i\partial) \psi_1$

$\psi_1 \rightarrow Z \psi_1$  wavefn renormalization. Likewise for scalars. Entire superfield renormalizes as

$\overline{\Phi}_1 \rightarrow Z \overline{\Phi}_1$  i.e.  $\int d^4\theta \overline{\Phi}_1 \Phi_1 \rightarrow \int d^4\theta Z^2 \overline{\Phi}_1 \Phi_1$

evaluate above diag  $\sim \overline{g} g \int d^4k \frac{1}{(k-m)} \frac{1}{((p-k)^2 - m^2)}$

○ extract term  $\sim \not{p}$  so  $\int \frac{d^4k}{k^4} \log$  diverg.

get  $Z = 1 + \overset{\text{tree}}{\downarrow} g\bar{g} \overset{\text{1 loop}}{\downarrow} \ln \left| \frac{\mu}{m} \right|^2 + \dots$

with  $m = \langle \phi_1 \rangle$

so  $K_{\text{eff}} = \overline{\Phi}_1 \Phi_1 \left( 1 + \overset{\text{tree}}{\swarrow} g\bar{g} \overset{\text{1 loop}}{\swarrow} \ln \left( \frac{\mu^2}{\overline{\Phi}_1 \Phi_1} \right) + \dots \right)$  higher order  
↓

General feature

$$\mathcal{L}_{cl} = \int d^4\theta K_{cl} + \int d^2\theta W_{cl} + \text{h.c.}$$

quantum effective action

$$\mathcal{L}_{\text{eff}} = \int d^4\theta K_{\text{eff}} + \int d^2\theta W_{\text{eff}} + \text{h.c.}$$

generally  $K_{\text{eff}}$  gets complicated corrections in perturbation theory.

But generally  $W_{\text{eff}}$  gets no renormalization

in perturbation theory. Only possible

$W$  renormalization is via non-perturbative effects