## 4/29/13 Lecture outline

- $\star$  Reading: Zwiebach chapters 5, 6
- Last time: string action:  $S = \int dt dx \mathcal{L}(\frac{\partial y}{\partial t}, \frac{\partial y}{\partial x})$ , with momentum densities

$$\mathcal{P}^t = \frac{\partial \mathcal{L}}{\partial \dot{y}}, \qquad \mathcal{P}^x = \frac{\partial \mathcal{L}}{\partial y'}$$

Least action gives the equations of motion

$$\frac{\partial \mathcal{P}^t}{\partial t} + \frac{\partial \mathcal{P}^x}{\partial x} = 0$$

The non-relativistic string has  $\mathcal{L} = \frac{1}{2}\mu_0 \dot{y}^2 - \frac{1}{2}T_0 y'^2$ , which we're going to replace with a relativistic version. For guidance, noted that a relativistic point particle of mass mhas  $S = -mc \int ds = -mc^2 \int dt \sqrt{1 - v^2/c^2}$  and noted its reparametrization invariance: write  $x_\mu(\tau)$ , and can change worldline parameter  $\tau$  to an arbitrary new parameterization  $\tau'(\tau)$ , and the action is invariant. To see this use  $S = -mc \int \sqrt{-\eta_{\mu\nu} \frac{dx^{\mu}}{d\tau} \frac{dx^{\nu}}{d\tau}}$  and change  $\frac{dx^{\mu}}{d\tau} = \frac{dx^{\mu}}{d\tau'} \frac{d\tau'}{d\tau}$  and note that  $S \to S$ . Euler Lagrange equations of motion:  $\frac{dp_{\mu}}{d\tau} = 0$ .

• When the particle is charged and in the presence of electric and magnetic fields, there is the new term in the action

$$S = \int (-mcds + \frac{q}{c}A_{\mu}dx^{\mu}), \qquad (1)$$

which is manifestly relativistically invariant (and also repparameterization) invariant. Note also that, under a gauge transformation, we have  $S \to S + \frac{qf}{c}$ , which does not affect the equations of motion (just as changing the Lagrangian by a total time derivative does not).

The lagrangian is thus  $L = -mc\sqrt{1 - \vec{v}^2/c^2} + \frac{q}{c}\vec{v}\cdot\vec{A} - q\phi$ . The momentum conjugate to  $\vec{r}$  is  $\vec{P} = \partial L/\partial \vec{v} = m\vec{v}/\sqrt{1 - \vec{v}^2/c^2} + \frac{q}{c}\vec{A}$ . The Hamiltonian is  $H = \vec{v}\cdot\vec{P} - L = \sqrt{m^2c^4 + c^2(\vec{P} - \frac{q}{c}\vec{A})^2} + q\phi$ . The equations of motion can be written as  $\frac{d^2x^{\mu}}{d\tau^2} = \frac{q}{mc}F_{\mu\nu}\frac{dx^{\nu}}{d\tau}$ . In the non-relativistic limit we have  $H = \frac{1}{2m}(\vec{P} - \frac{q}{c}\vec{A})^2 + q\phi$ , where  $\vec{P} - \frac{q}{c}\vec{A} = m\vec{v}$ .

• Recap:  $S = -mc \int ds + \frac{q}{c} \int A_{\mu} dx^{\mu}$  for a relativistic point particle, where we can write  $ds = \sqrt{-g_{\mu\nu}\dot{x}^{\mu}\dot{x}^{\nu}}d\tau$ , with  $\dot{=} \frac{d}{d\tau}$ , and  $\tau$  is the arbitrary worldline parameter, with reparameterization symmetry  $\tau \to \tau'$ . For a string world-sheet, we need two parameters,  $\xi^a$ , a = 1, 2. The string trajectory is  $x : \Sigma \to M$ , where  $\Sigma$  is the 2d world-sheet, with local coordinates  $\xi^a$ , and M is the target space, with local coordinates  $x^{\mu}$ . The worldsheet area element is  $A = \int d^2\xi \sqrt{|h|}$ , where  $h_{ab}$  is the worldsheet metric, and |h| is its determinant.

Suppose that the target space has metric  $g_{\mu\nu}$ , with space-time length e.g.  $ds^2 = g_{\mu\nu}dx^{\mu}dx^{\nu}$ . By writing  $dx^{\mu} = \partial_a x^{\mu}d\xi^a$ , we get

$$ds^{2} = g_{\mu\nu} \frac{dx^{\mu}}{d\xi^{a}} \frac{dx^{\nu}}{d\xi^{b}} d\xi^{a} d\xi^{b}, \qquad \text{so} \qquad h_{ab} = g_{\mu\nu} \frac{dx^{\mu}}{d\xi^{a}} \frac{dx^{\nu}}{d\xi^{b}},$$

where this  $h_{ab}$  is called the induced metric. So the worldsheet area functional is

$$A = \int d^2 \xi \sqrt{\det(g_{\mu\nu} \frac{dx^{\mu}}{d\xi^a} \frac{dx^{\nu}}{d\xi^b})}.$$