## 4/29/13 Lecture outline

- $\star$  Reading: Zwiebach chapters 5, 6
- Last time: string action:  $S = \int dt dx \mathcal{L}(\frac{\partial y}{\partial t}, \frac{\partial y}{\partial x})$ , with momentum densities

$$
\mathcal{P}^t = \frac{\partial \mathcal{L}}{\partial \dot{y}}, \qquad \mathcal{P}^x = \frac{\partial \mathcal{L}}{\partial y'}.
$$

Least action gives the equations of motion

$$
\frac{\partial \mathcal{P}^t}{\partial t} + \frac{\partial \mathcal{P}^x}{\partial x} = 0.
$$

The non-relativistic string has  $\mathcal{L} = \frac{1}{2}$  $\frac{1}{2}\mu_0\dot{y}^2-\frac{1}{2}$  $\frac{1}{2}T_0y^2$ , which we're going to replace with a relativistic version. For guidance, noted that a relativistic point particle of mass  $m$ has  $S = -mc \int ds = -mc^2 \int dt \sqrt{1 - v^2/c^2}$  and noted its reparametrization invariance: write  $x_{\mu}(\tau)$ , and can change worldline parameter  $\tau$  to an arbitrary new parameterization  $\tau'(\tau)$ , and the action is invariant. To see this use  $S = -mc \int \sqrt{-\eta_{\mu\nu} \frac{dx^{\mu}}{d\tau}}$  $d\tau$  $\frac{dx^{\nu}}{d\tau}$  and change  $\frac{dx^{\mu}}{d\tau} = \frac{dx^{\mu}}{d\tau'}$  $d\tau'$  $\frac{d\tau'}{d\tau}$  and note that  $S \to S$ . Euler Lagrange equations of motion:  $\frac{dp_{\mu}}{d\tau} = 0$ .

• When the particle is charged and in the presence of electric and magnetic fields, there is the new term in the action

$$
S = \int (-mcds + \frac{q}{c}A_{\mu}dx^{\mu}), \qquad (1)
$$

which is manifestly relativistically invariant (and also repparameterization) invariant. Note also that, under a gauge transformation, we have  $S \to S + \frac{qf}{c}$  $\frac{dI}{c}$ , which does not affect the equations of motion (just as changing the Lagrangian by a total time derivative does not).

The lagrangian is thus  $L = -mc\sqrt{1 - \vec{v}^2/c^2} + \frac{q}{c}$  $\frac{q}{c}\vec{v}\cdot\vec{A}-q\phi$ . The momentum conjugate to  $\vec{r}$  is  $\vec{P} = \partial L/\partial \vec{v} = m\vec{v}/\sqrt{1-\vec{v}^2/c^2} + \frac{q}{c}\vec{A}$ . The Hamiltonian is  $H = \vec{v} \cdot \vec{P} - L =$  $\sqrt{m^2c^4+c^2(\vec{P}-\frac{q}{c}\vec{A})^2}+q\phi$ . The equations of motion can be written as  $\frac{d^2x^{\mu}}{d\tau^2}=\frac{q}{m}$  $\frac{q}{mc}F_{\mu\nu}\frac{dx^{\nu}}{d\tau}.$ In the non-relativistic limit we have  $H = \frac{1}{2m}(\vec{P} - \frac{q}{c}\vec{A})^2 + q\phi$ , where  $\vec{P} - \frac{q}{c}\vec{A} = m\vec{v}$ .

• Recap:  $S = -mc \int ds + \frac{q}{c}$  $\frac{q}{c} \int A_{\mu} dx^{\mu}$  for a relativistic point particle, where we can write  $ds = \sqrt{-g_{\mu\nu}\dot{x}^{\mu}\dot{x}^{\nu}}d\tau$ , with  $\dot{=} \frac{d}{d\tau}$ , and  $\tau$  is the arbitrary worldline parameter, with reparameterization symmetry  $\tau \to \tau'$ . For a string world-sheet, we need two parameters,  $\xi^a$ ,  $a = 1, 2$ . The string trajectory is  $x : \Sigma \to M$ , where  $\Sigma$  is the 2d world-sheet, with local coordinates  $\xi^a$ , and M is the target space, with local coordinates  $x^{\mu}$ . The worldsheet area element is  $A = \int d^2 \xi \sqrt{|h|}$ , where  $h_{ab}$  is the worldsheet metric, and |h| is its determinant.

Suppose that the target space has metric  $g_{\mu\nu}$ , with space-time length e.g.  $ds^2 = g_{\mu\nu} dx^{\mu} dx^{\nu}$ . By writing  $dx^{\mu} = \partial_a x^{\mu} d\xi^a$ , we get

$$
ds^2 = g_{\mu\nu}\frac{dx^{\mu}}{d\xi^a}\frac{dx^{\nu}}{d\xi^b}d\xi^a d\xi^b, \qquad \text{so} \qquad h_{ab} = g_{\mu\nu}\frac{dx^{\mu}}{d\xi^a}\frac{dx^{\nu}}{d\xi^b},
$$

where this  $h_{ab}$  is called the induced metric. So the worldsheet area functional is

$$
A = \int d^2 \xi \sqrt{\det(g_{\mu\nu} \frac{dx^{\mu}}{d\xi^{a}} \frac{dx^{\nu}}{d\xi^{b})}}.
$$