4/17/09 Lecture outline

 \star Reading: Zwiebach chapter 4 and 5.

• Recall that $[S] = ML^2/T$, same as [\hbar]. (Indeed, Feynman's formulation of QM is based on $\psi \sim \int_{\text{all paths } x(t)} [dx(t)] e^{iS[x(t)]/\hbar}$.

As mentioned last time, the action for a relativistic point particle of mass m is $S =$ $-mc \int ds = -mc^2 \int dt \sqrt{1 - v^2/c^2}$. This gives $\vec{p} = \partial_{\vec{v}} = \gamma m \vec{v}$ and $H = \vec{p} \cdot \vec{v} - L = \gamma mc^2$, both of which are constants of the motion (thanks to the time and spatial translation invariance).

• Reparametrization invariance: write $x_{\mu}(\tau)$, and can change worldline parameter τ to an arbitrary new parameterization $\tau'(\tau)$, and the action is invariant. To see this use $S = -mc \int \sqrt{-\eta_{\mu\nu} \frac{dx^{\mu}}{d\tau}}$ $d\tau$ $\frac{dx^{\nu}}{d\tau}$ and change $\frac{dx^{\mu}}{d\tau} = \frac{dx^{\mu}}{d\tau'}$ $d\tau'$ $\frac{d\tau'}{d\tau}$ and note that $S \to S$. The Euler Lagrange equations of motion are $\frac{dp_{\mu}}{d\tau} = 0$.

When the particle is charged and in the presence of electric and magnetic fields, there is the new term in the action $S = -mc \int ds + \frac{q}{c}$ $\frac{q}{c} \int A_{\mu} dx^{\mu}$, which is also reparameterization invariant. The equations of motion can now be written as $\frac{d^2x^{\mu}}{d\tau^2} = \frac{q}{m}$ $\frac{q}{mc}F_{\mu\nu}\frac{dx^{\nu}}{d\tau}.$

• A key concept from last lecture is that E&M is associated with the local gauge transformation

$$
\psi(t,\vec{x}) \to e^{iqf(t,\vec{x})/\hbar c} \psi(t,\vec{x}), \qquad A_{\mu} \to A_{\mu} + \partial_{\mu} f(t,\vec{x}). \tag{1}
$$

According to Noether's theorem, there is a one-to-one correspondence

(continuous) global symmetry \leftrightarrow conserved quantity.

The original example the relation between translation symmetry in time and/or space, $x^{\mu} \rightarrow x^{\mu} + a^{\mu}$, and conservation of energy and/or momentum, p^{μ} .

There is another deep correspondence

$$
local gauge symmetry \rightarrow forces.
$$

and E&M is the force associated with the local symmetry above. There is still a conserved charge, in E&M it is current conservation $\partial^{\mu} j_{\mu} = 0$. As we'll now discuss, in general relativity (GR) the above spacetime translation symmetry is a subgroup of a more general symmetry, general coordinate invariance, which is the fundamental symmetry principle associated with gravity.

• A brief (!) introduction to general relativity. We replace the metric $\eta_{\mu\nu}$ with a dynamical quantity $g_{\mu\nu}$. There is a symmetry principle which is akin to the gauge invariance of electricity and magnetism and to the above reprarameterization invariance. It is general coordinate invariance: $x^{\mu} \to x^{\mu'}(x^{\mu})$. Physics is invariant under such local coordinate changes. The metric transforms as $g_{\mu\nu} = g_{\mu'\nu'} \frac{dx^{\mu'}}{dx^{\mu}}$ dx^{μ} $\frac{dx^{\nu'}}{dx^{\nu}}$. The action of a point particle is $S = -mc \int ds + \frac{q}{c}$ $\frac{q}{c} \int A_{\mu} dx^{\mu}$, just like before, except that we contract and raise and lower indices with $g_{\mu\nu}$ rather than $\eta^{\mu\nu}$. Get from the Euler Lagrange equations now

$$
\frac{d^2x^{\mu}}{d\tau^2} + \Gamma^{\mu}_{\rho\sigma} \frac{dx^{\rho}}{d\tau} \frac{dx^{\sigma}}{d\tau} = \frac{q}{mc} F^{\mu}_{\nu} \frac{dx^{\nu}}{d\tau},
$$

where

$$
\Gamma^{\mu}_{\rho\sigma} = \frac{1}{2}g^{\mu\lambda}(\partial_{\rho}g_{\lambda\sigma} + \partial_{\sigma}g_{\lambda\rho} - \partial_{\lambda}g_{\rho\sigma})
$$

is the connection; it is analogous to A^{μ} in electromagnetism. The connection enters into covariant derivatives like $\nabla_{\rho}V^{\mu} = \partial_{\rho}V^{\mu} + \Gamma^{\mu}_{\rho\sigma}V^{\sigma}$ in order to have things transform properly under general coordinate transformations (analogous to the gauge invariant covariant derivatives $D_{\mu} = \partial_{\mu} - i \frac{q}{\hbar c} A_{\mu}$ in E&M). The above equations of motion is called the geodesic equation; it reparameterization invariant $(\tau \to \tau')$ and transforms properly under general coordinate transformations $x^{\mu} \rightarrow x^{\mu'}$.

The Riemann tensor is

$$
R^{\rho}_{\sigma\mu\nu} = \partial_{\mu}\Gamma^{\rho}_{\nu\sigma} + \Gamma^{\rho}_{\mu\lambda}\Gamma^{\lambda}_{\nu\sigma} - (\mu \leftrightarrow \nu).
$$

It is analogous to $F_{\mu\nu}$ in E&M. The Ricci tensor is $R_{\mu\nu} = R^{\lambda}_{\mu\lambda\nu}$ and the Ricci scalar is $R = R^{\mu}_{\mu}$. The metric is dynamically determined by minimizing the action w.r.t. $\delta g_{\mu\nu}$, where there is a term

$$
S = \frac{1}{16\pi G_D} \int d^D x \sqrt{|g|} R + \dots
$$

For fun, we wrote it in general spacetime dimension D. Let's note the units (setting $c=1$): $[R] = L^{-2}$ and $[S] = ML$, so $[G_D] = L^{D-3}M^{-1}$. Since $[\hbar] = ML$, we have $G_D = \ell_P^{D-2}$ P in D spacetime dimensions. (Note that $\int d^D x \sqrt{|g|}$ gives the spacetime volume (which is clearly general coordinate invariant). This comment will be useful very soon, when we write down the relativistic string action!)

Note also that the relation $G_D = GV_C$ is evident from the above action.

In the weak curvature limit, we can reduce to the gravitational potentials, with $\nabla^2 V_g^{(D)} = 4\pi G_D \rho_m$. This comes from $g_{\mu\nu} \approx \eta_{\mu\nu} + h_{\mu\nu}$ and $h_{0,0} \approx -2V_g$.

• Nonrelativistic strings. $[T_0] = [F] = [E]/L = [\mu_0][v^2]$. Indeed, considering $F = ma$ for an element dx of the string yields the string wave equation $\frac{\partial^2 y}{\partial x^2} - \frac{1}{v_c^2}$ v_0^2 $\frac{\partial^2 y}{\partial t^2} = 0$, with $v_0 = \sqrt{T_0/\mu_0}$. Endpoints at $x = 0$ and $x = a$. Can choose Dirichlet or Neumann BCs at these points. With Dirichlet at each end, $y_n(x) = A_n \sin(n\pi x/a)$ and the general solution is $y(x,t) = \sum_n y_n(x) \cos \omega_n t$, where $\omega_n = v_0 n \pi/a$ (and the A_n are determined from the initial conditions, by Fourier transform).

The nonrelativistic string action is $S = \int dt L$ where L is the kinetic energy minus potential energy, which gives

$$
S = \int dt \int dx \left(\frac{1}{2}\mu_0 \left(\frac{\partial y}{\partial t}\right)^2 - \frac{1}{2}T_0 \left(\frac{\partial y}{\partial x}\right)^2 \right),
$$

which is a particular case of the more general action $S = \int dt dx \mathcal{L}(\frac{\partial y}{\partial t}, \frac{\partial y}{\partial x})$. We can then define the momentum density and corresponding quantity

$$
\mathcal{P}^t = \frac{\partial \mathcal{L}}{\partial \dot{y}}, \qquad \mathcal{P}^x = \frac{\partial \mathcal{L}}{\partial y'},
$$

and the action is made stationary, $\delta S = 0$, if

$$
\frac{\partial \mathcal{P}^t}{\partial t} + \frac{\partial \mathcal{P}^x}{\partial x} = 0,
$$

which when applied to the above particular choice of action gives the usual wave equation. Note that Neumann or Dirichlet BCs correspond to $\mathcal{P}^x = 0$ or $\mathcal{P}^t = 0$ at the boundary, respectively, and that this is indeed needed for the surface terms to be compatible with $\delta S = 0.$