

10/24/16 Lecture 9 outline

• Last time: position space probability density  $\rho(\vec{x}, t) = |\psi(\vec{x}, t)|^2$  and current  $\vec{j}(\vec{x}, t) = (\hbar/m)\text{Im}(\psi^* \nabla \psi)$ . Note  $\int d^3\vec{x} \vec{j} = \langle \vec{p} \rangle / m$ . Can also write  $\vec{j} = \rho \nabla S / m$ , where  $\psi \equiv \sqrt{\rho} e^{iS/\hbar}$ . E.g. for a plane wave  $\nabla S = \vec{p}$ . Substituting  $\psi \equiv \sqrt{\rho} e^{iS/\hbar}$  into the time dependent SE gives an equation where each  $S$  derivative has a  $1/\hbar$ . In the classical limit we have e.g.  $|\nabla S|^2 \gg \hbar |\nabla^2 S|$  and the SE reduces to

$$\frac{1}{2m} |\nabla S|^2 + V(x) + \frac{\partial S(\vec{x}, t)}{\partial t} = 0$$

which is the Hamilton-Jacobi equation of classical mechanics with  $S$  Hamilton's function. This shows how the SE reduces to classical mechanics in the  $S/\hbar \ll 1$  limit. We will soon briefly discuss the path integral description of QM, where  $S$  is replaced with the action functional.

For an energy eigenstate,  $S = W(x) - Et$ .

Example: 3d harmonic oscillator, ground state is  $|000\rangle$  and  $\psi_{000}(\vec{x}) = c_0^3 \psi_0(x) \psi_0(y) \psi_0(z) = c_0^3 e^{-m\omega r^2/2\hbar}$ . First excited states are  $|100\rangle, |010\rangle, |001\rangle$ , 3-fold degenerate. We'll soon discuss spherically symmetric potentials more generally and see that this degeneracy is related to angular momentum and spherical harmonics,  $\ell = 1, m = 1, 0, -1$ . For now just mention it and illustrate  $\rho$  and  $\vec{j}$ .

• Linear potential,  $V = k|x|$ . This case does not have a simple solution in terms of a trick – one has to consider the differential equation. Physically it is also less interesting than the SHO. It comes up e.g. for a particle in a constant force field (e.g. a gravitational field close to the earth's surface), if we replace the  $x < \infty$  potential with an infinite one. The  $|x|$  is not nice near  $x = 0$ ; not physically realistic there. The main reason to mention it is because an arbitrary potential, in the vicinity of its turning point, can be approximated as a linear potential by the first term in the Taylor expansion. The SE can be converted by a change of variables into  $u_E''(z) - zu_E(z) = 0$ , which is the Airy equation, whose solution is  $Ai(z)$ . It oscillates for  $z < 0$  and has exponential decay for  $z > 0$ , which is the classically forbidden region.  $Ai(z) \rightarrow z^{-1/4} (2\sqrt{\pi})^{-1} e^{-2z^{3/2}/3}$  for  $z \rightarrow \infty$  and  $Ai(z) \rightarrow |z|^{-1/4} \pi^{-1/2} \cos(2/3|z|^{3/2} - \pi/4)$  for  $z \rightarrow -\infty$ .

The energy levels are determined by the condition that either  $u_E$  or  $u_E'$  vanishes at  $x = 0$  (parity) so  $E$  is quantized according to the zeros of  $Ai$  or  $Ai'$ .