10/5 Lecture outline

\star Reading for today's lecture: Coleman to end of lecture 4 (p. 37).

• Last time: symmetries of $\mathcal L$ and Noether's theorem. If a variation $\delta\phi_a$ changes $\delta \mathcal{L} = \partial_{\mu} F^{\mu}$, then it's a symmetry of the action and there is a conserved current: $j^{\mu} =$ ∂L $\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_a)} \delta \phi_a - F^\mu.$

Translation invariance: $x^{\mu} \to x^{\mu} + \epsilon^{\mu}$, $\delta \phi_a = \epsilon^{\nu} \partial_{\nu} \phi_a$, $\delta \mathcal{L} = \epsilon^{\nu} \partial_{\nu} \mathcal{L}$ (assuming no explicit x dependence). Get $T_{\mu\nu} = \frac{\partial \mathcal{L}}{\partial \partial \mu}$ $\frac{\partial \mathcal{L}}{\partial \partial_{\mu}\phi_{a}}\partial_{\nu}\phi_{a} - g_{\mu\nu}\mathcal{L}$. Stress energy tensor. Conserved charge is $P_{\mu} = \int d^3 \vec{x} T_{\mu 0}$. Another example: $x^{\mu'} = \Lambda^{\mu'}_{\nu} x^{\nu}$ Lorentz boost and rotation symmetry leads to conservation of angular momentum. Write $\Lambda^{\mu}_{\nu} = \delta^{\mu}_{\nu} + \omega^{\mu}_{\nu}$, leads to conserved $M_{\mu\rho\sigma} = x_{\mu}T_{\rho\sigma} - x_{\sigma}T_{\rho\mu}$. Conserved charge is $M_{\rho\sigma} = \int d^3x M_{0\rho\sigma}$.

• Apply to $\mathcal{L}_{KG} = \frac{1}{2}$ $\frac{1}{2}\partial_{\mu}\phi\partial^{\mu}\phi - \frac{1}{2}m^2\phi^2$. Get $T_{\mu\nu} = \partial_{\mu}\phi\partial_{\nu}\phi - \frac{1}{2}$ $\frac{1}{2}\eta_{\mu\nu}\partial_{\lambda}\phi\partial^{\lambda}\phi + \frac{1}{2}m^2\phi^2\eta_{\mu\nu}$. So

$$
H = \int d^3x \mathcal{H}, \qquad \vec{P} = \int d^3x \vec{\mathcal{P}}.
$$

$$
\mathcal{H} = \frac{1}{2}\dot{\phi}^2 + \frac{1}{2}(\nabla\phi)^2 + \frac{1}{2}m^2\phi^2, \qquad \vec{\mathcal{P}} = \dot{\phi}\nabla\phi.
$$

• Recall from last week: $SHO = KG$ equation in $0 + 1$ dimensions, i.e. the SHO: $L=\frac{1}{2}$ $rac{1}{2}\dot{\phi}^2 - \frac{1}{2}$ $\frac{1}{2}\omega^2\phi^2$, $\Pi = \partial L/\partial \dot{\phi} = \dot{\phi}$. Now quantize: $[\phi, \Pi] = i\hbar$, $[a, a^{\dagger}] = 1$, $H = \omega(a^{\dagger}a + \frac{1}{2})$ $(\frac{1}{2})$. Heisenberg picture, $\widehat{\phi} = \sqrt{\frac{1}{2\omega}}$ $\frac{1}{2\omega}(ae^{-i\omega t} + a^{\dagger}e^{i\omega t}); \Pi = \dot{\phi} = -i\sqrt{\frac{\omega}{2}}(ae^{i\omega t} - a^{\dagger}e^{-i\omega t}).$ Define $|0\rangle$ s.t. $a|0\rangle = 0$, and $|n\rangle = c_n(a^{\dagger})^n|0\rangle$.

• Canonical quantization: generalize QM by replacing $q_a(t) \to \phi(t, \vec{x})$. It's conjugate momentum is $\Pi \equiv \partial \mathcal{L}/\partial \dot{\phi}$. The theory is quantized by replacing ϕ and Π with operators (sometimes we'll give them hats, but usually won't bother), satisfying

$$
[\phi_a(\vec{x},t), \Pi_b(\vec{y},t)] = i\delta_{ab}\delta^3(\vec{x}-\vec{y})
$$
 (Equal time commutators).

$$
[\phi_a(\vec{x},t), \phi_b(\vec{y},t)] = 0.
$$

• Quantize the KG field theory in $3 + 1$ dimensions. Write

$$
\phi(x) = \int \frac{d^3k}{(2\pi)^3} \frac{1}{\sqrt{2\omega_{\vec{k}}}} [a_{\vec{k}}e^{-ikx} + a_{\vec{k}}^{\dagger}e^{ikx}],
$$

$$
\Pi(x) = \dot{\phi}(x) = \int \frac{d^3k}{(2\pi)^3} (-i)\sqrt{\frac{\omega_{\vec{k}}}{2}} [a_{\vec{k}}e^{-ikx} - a_{\vec{k}}^{\dagger}e^{ikx}],
$$

Then canonical quantization implies that

$$
[a_{\vec{k}},a^\dagger_{\vec{k}'}] = (2\pi)^3 \delta^3(\vec{k}-\vec{k}'),
$$

i.e. they're creation and annihilation operators, with others vanishing. It will be useful to define $a(k) \equiv \sqrt{2\omega_k} a_{\vec{k}}$, so then the above becomes

$$
\phi(x) = \int \frac{d^3k}{(2\pi)^3 2\omega(k)} [a(k)e^{-ikx} + a^{\dagger}(k)e^{ikx}],
$$

$$
[a(k), a^{\dagger}(k')] = (2\pi)^3 2\omega_k \delta^3(\vec{k} - \vec{k}'),
$$

with the relativistic-invariant measures appearing.

The quantum field ϕ is a superposition of creation and annihilation operators. Also, plugging into our expressions for energy and momentum gives the operators

$$
H = \frac{1}{2} \int \frac{d^3k}{(2\pi)^2 (2\omega)} \omega(a(\vec{k})a^{\dagger}(\vec{k}) + a^{\dagger}(\vec{k})a(\vec{k})),
$$

$$
\vec{P} = \frac{1}{2} \int \frac{d^3k}{(2\pi)^2 (2\omega)} \vec{k}(a(\vec{k})a^{\dagger}(\vec{k}) + a^{\dagger}(\vec{k})a(\vec{k})),
$$

Need to normal order the first term. Define : AB : for operators A and B to mean that the terms are arranged so that the annihilation operators are on the right, so annihilates the vacuum.

• The vacuum $|0\rangle$ is annihilated by all $a(k)$. Create states with momenta p_1^{μ} $j_1^{\mu}, \ldots, p_n^{\mu}$ via $a^{\dagger}(p_1) \dots a^{\dagger}(p_n)|0\rangle$. Note that these behave as identical bosons: the state is symmetric under exchanging any pair of momenta, because $[a^{\dagger}(p), a^{\dagger}(p')] = 0$.

• Two-point field correlation function:

$$
\langle 0|\phi(x)\phi(y)|0\rangle \equiv D(x-y) = \int \frac{d^3k}{(2\pi)^3 2\omega(k)} e^{-ik(x-y)}.
$$

Note also that $2i\partial_{x^0}D(x - y)$ is the integral that we saw in last lecture, for the probability amplitude to find a particle having traveled with spacetime displacement $(x - y)^{\mu}$. For spacelike separation, $(x-y)^2 = -r^2$, we here get $D(x-y) = \frac{m}{2\pi^2 r} K_1(mr)$, with K_1 a Bessel function. Recall that the Bessel function has a simple pole when its argument vanishes, and exponentially decays at infinity. So $D(x - y) \sim \exp(-m|\vec{x} - \vec{y}|)$ is non-vanishing outside the forward light cone.