## 10/31 Lecture outline

- \* Reading for today's lecture: Luke p. 65-80; Tong p. 35-41.
- Last time: amplitudes in our nucleon + meson toy model, via

$$\langle f|(S-1)|i\rangle = \langle f|Te^{-i\int d^4x \mathcal{H}_I(x)}|i\rangle \equiv i\mathcal{A}_{fi}(2\pi)^4 \delta^{(4)}(p_f - p_i).$$

Examples:  $N + N \to N + N$ , to  $\mathcal{O}(g^2)$ :

$$\mathcal{A} = (-ig)^2 \left[ \frac{1}{(p_1 - p_1')^2 - \mu^2} + \frac{1}{(p_1 - p_2')^2 - \mu^2} \right].$$

(1)  $N(p_1) + \bar{N}(p_2) \to N(p_1') + \bar{N}(p_2')$  has

$$i\mathcal{A} = (-ig)^2 \left( \frac{i}{(p_1 - p'_1) - \mu^2} + \frac{i}{(p_1 + p_2) - \mu^2} \right).$$

(2)  $N(p_1) + \bar{N}(p_2) \to \phi(p_1')\phi(p_2')$  has

$$i\mathcal{A} = (-ig)^2 \left( \frac{i}{(p_1 - p_1') - m^2} + \frac{i}{(p_1 - p_2') - m^2} \right).$$

(3)  $N(p_1) + \phi(p_2) \to N(p_1') + \phi(p_2')$  has

$$i\mathcal{A} = (-ig)^2 \left( \frac{i}{(p_1 - p_2') - m^2} + \frac{i}{(p_1 + p_2) - m^2} \right).$$

- Mandelstam variables.  $s = (p_1 + p_2)^2$ ,  $t = (p_1 p_1')^2$ ,  $u = (p_1 p_2')^2$ , with  $s + t + u = 4m^2$  (more generally,  $s + t + u = \sum_{i=1}^4 m_i^2$ ). In CM,  $s = 4E^2$ ,  $t = -2\vec{p}^2(1 \cos\theta)$ , and  $u = -2\vec{p}^2(1 + \cos\theta)$ .
- Crossing symmetry, CPT. Write  $1+2\to \bar 3+\bar 4$ . Take all momenta incoming,  $\mathcal{A}(p_1,p_2,p_3,p_4)$ , with  $p_1+p_2+p_3+p_4=0$  and use  $s=(p_1+p_2)^2,\ t=(p_1+p_3)^2$  and  $u=(p_1+p_4)^2$ . Note  $s+t+u=\sum_{n=1}^4 m_n^2$ . The process  $1+2\to \bar 3+\bar 4$  is kinematically allowed for  $s>4m^2,\ t<0,\ u<0$ . If instead  $u>4m^2$ , it's the process  $1+3\to \bar 2+\bar 4$ .
- Yukawa potential. Indeed, the t-channel term in e.g. the above N+N scattering amplitude gives, upon using  $(p_1-p_1')^2-\mu^2=-(|\vec{p_1}-\vec{p_1'}|^2+\mu^2)$ , and the Born approximation in NRQM,  $\mathcal{A}_{NR}=\int d^3\vec{r}e^{-i(\vec{p'}-\vec{p})}V(\vec{r})$ , the attractive Yukawa potential

$$V(r) = \int \frac{d^3q}{(2\pi)^3} \frac{-(g/2m)^2 e^{i\vec{q}\cdot\vec{r}}}{|\vec{q}|^2 + \mu^2} = -\frac{(g/2m)^2}{4\pi r} e^{-\mu r}.$$

<sup>1</sup> Max Born, in QM, or Lord Rayleigh classically:  $\frac{d\sigma}{d\Omega} \sim |U(\vec{q})|^2$ 

(The  $1/(2m)^2$  is because we normalized the relativistic states with the extra factor of  $2E \approx 2m$  as compared with standard nonrelativistic normalization<sup>2</sup>. This gives Yukawa's explanation of the attraction between nucleons, from meson exchange. The u-channel term is an exchange potential interaction, which exchanges the positions of the two identical particles in addition to giving a potential. For angular momentum  $\ell$  in a partial-wave expansion, the exchange term differs from the direct one by a factor of  $(-1)^{\ell}$ .

• We saw above that the t channel term above is associated with the Yukawa potential. The u channel term is similar. Now consider the s channel, in e.g. the  $N+\bar{N}$  scattering amplitude. Using the CM relations  $\vec{p}_1 = -\vec{p}_2 \equiv \vec{p}$  and  $E_1 = E_2 = \sqrt{p^2 + m^2}$  gives

$$\mathcal{A} \sim \frac{1}{4m^2 + 4p^2 - \mu^2 + i\epsilon},$$

so for  $\mu < 2m$  the denominator is always positive, and the amplitude decreases with increasing  $p^2$ . For  $\mu > 2m$  there is a pole at  $(p_1 + p_2)^2 = \mu^2$ , where the intermediate meson goes on shell. This leads to a peak (not a pole, of course; because the intermediate particle is unstable anyway, the denominator gets an imaginary contribution from higher order contributions), a resonance, in the cross section. E.g.  $Z_0$  pole in  $e^+e^- \to \mu^+\mu^-$ , but not in  $e^+e^- \to \gamma\gamma$ .

- Solve  $\mathcal{L} = \frac{1}{2}\partial\phi^2 \frac{1}{2}m^2\phi^2 J(x)\phi$ . Using Dyson + Wick's theorem,  $U(\infty, -\infty) =:$   $e^{O_1 + \frac{1}{2}O_2}:$ , where  $O_1 = -i\int d^4x J(x)\phi(x)$  and  $O_2 = (-i)^2\int d^4x_1d^4x_2D_F(x_1-x_2)J(x_1)J(x_2)$ . So  $O_2 = \alpha + i\beta$  is a number, whereas  $O_1$  is an operator. Will lead to probability  $P_n$  for creating out of the vacuum a state with n mesons given by  $P_n = e^{-|\alpha|}|\alpha|^n/n!$ , the Poisson distribution. You'll work out the details in the HW assignment.
- Compute probabilities by squaring the S-maxtrix amplitudes, but have to be careful with the delta functions, since squaring the delta functions would give nonsense.

Warmup: consider quantum mechanics, with  $U(t) = Te^{-i\int_{0}^{t} H(t)dt}$ ,

$$\langle f|U(t)|i\rangle \approx -i\langle f|H_{int}|i\rangle \int_0^t dt e^{i\omega t},$$

where  $\omega = E_f - E_i$ . If we take  $t \to \infty$  first, we get  $\delta(\omega)$  and squaring would give nonsense. That's because we're asking the wrong question if we ask about probability for a transition over all time – instead, we should ask about the rate. So keep t finite for now.

<sup>&</sup>lt;sup>2</sup> This is clear on dimensional grounds, since  $[g] \sim m$ . More generally, write  $a(p) = \sqrt{2E}\widehat{a}(p)$  and  $\mathcal{A} = \prod_i \sqrt{2E_i} \prod_f \sqrt{2E_f} \widehat{\mathcal{A}}$ .

Squaring gives  $P(t) = 2|\langle f|H_{int}|i\rangle|^2(1-\cos\omega t)/\omega^2$ . For  $t\to\infty$ , multiply by  $dE_f\rho(E_f)$  and replace  $(1-\cos\omega t)/\omega^2 = 4\sin^2(\frac{1}{2}\omega t)/\omega^2 \to \pi t\delta(\omega)$  (using  $\int_{-\infty}^{\infty} dx x^{-2}\sin^2 x = \pi$  (hint:  $\sin^2 x/x^2 = (2-e^{i2x}-e^{-i2x})/4x^2$  and close the contour in the correct directions)) to get

$$\dot{P}_{i \to f} = 2\pi |\langle f | H_{int} | i \rangle|^2 \rho(E).$$

This is "Fermi's Golden Rule" – it was actually derived by Dirac, but Fermi used it a lot and called it the golden rule. Another aside: Fermi and Dirac independently discovered that spin 1/2 objects must anticommute, and Dirac generously named such objects "Fermions".

Naively taking  $t \to \infty$  initially would have given amplitude  $\sim \delta(\omega)$ , and squaring that would give  $\delta(\omega)^2$ , which needs to be replaced with  $\delta(\omega)2\pi T$ , and then divide by T to get the rate. Similarly in field theory,  $\delta(p)^2$  should be replaced with probability  $\sim \delta(p)$  times phase space volume factors.