

9/30 Lecture outline

★ **Reading: Luke, chapters 3 and 4. Maybe a bit of chapter 5, if time.**

• Last time: classical field theory. E.g. scalars $\phi_a(t, \vec{x})$, with $S = \int d^4x \mathcal{L}(\phi_a, \partial_\mu \phi_a)$. Then $\Pi_a^\mu = \partial \mathcal{L} / \partial (\partial_\mu \phi_a)$, and E.L. eqns $\partial \mathcal{L} / \partial \phi_a = \partial_\mu \Pi_a^\mu$. Define $\Pi_a \equiv \Pi_a^0$. $H = \int d^3x (\Pi_a \dot{\phi}_a - \mathcal{L}) = \int d^3x \mathcal{H}$.

Example: $\mathcal{L} = \frac{1}{2}(\partial_\mu \phi \partial^\mu \phi - m^2 \phi^2)$. Klein-Gordon equation, $(\partial^2 + m^2)\phi = 0$. Can't interpret ϕ as a probability wavefunction because of solutions $E = \pm \sqrt{\vec{p}^2 + m^2}$. But we'll see that the KG equation is fine as a field theory. The field has both creation and annihilation operators, corresponding to the $E = \pm \sqrt{\vec{k}^2 + m^2}$ solutions. Write general classical solution

$$\phi_{cl}(x) = \int \frac{d^3k}{(2\pi)^3(2\omega(k))} [a_{cl}(k)e^{-ikx} + a_{cl}^*(k)e^{ikx}],$$

where $a_{cl}(k)$ are classical constants of integration, determined by the initial conditions. We'll quantize soon.

Another example: $\mathcal{L} = \frac{i}{2}(\psi^* \dot{\psi} - \dot{\psi}^* \psi) - \nabla \psi^* \cdot \nabla \psi - m\psi^* \psi$. Get EOM: $i\partial_t \psi = -\nabla^2 \psi + m\psi$. Looks like S.E., but again don't want to interpret ψ as a probability amplitude – here it's a field, that we can consider quantizing. This example won't work for ψ a scalar field, but we'll later consider an analogous theory where ψ is a fermion field, and the equation is the Dirac equation.

• Canonical quantization: generalize QM by replacing $q_a(t) \rightarrow \phi_a(t, \vec{x})$. QM is like QFT in zero spatial dimensions, with the field playing role of position before:

$$[\phi_a(\vec{x}, t), \Pi_b(\vec{y}, t)] = i\delta_{ab}\delta^3(\vec{x} - \vec{y}) \quad (\text{Equal time commutators}).$$

• Consider the KG equation in 0 + 1 dimensions, i.e. the SHO: $L = \frac{1}{2}\dot{\phi}^2 - \frac{1}{2}\omega^2 x^2$, $p = \partial L / \partial \dot{\phi} = \dot{\phi}$. Classical EOM solved by $x_{cl} = ae^{-i\omega t} + a^*e^{i\omega t}$. Now quantize: $[x, p] = i\hbar$, $[a, a^\dagger] = 1$, $H = \omega(a^\dagger a + \frac{1}{2})$. In the Heisenberg picture, $\hat{x} = \sqrt{\frac{1}{2\omega}}(ae^{-i\omega t} + a^\dagger e^{i\omega t})$; $p = \dot{x} = i\sqrt{\frac{\omega}{2}}(ae^{i\omega t} - a^\dagger e^{-i\omega t})$.

• Now quantize the KG field theory in 3 + 1 dimensions. Write

$$\phi(x) = \int \frac{d^3k}{(2\pi)^3(2\omega(k))} [a(k)e^{-ikx} + a^\dagger(k)e^{ikx}].$$

Then canonical quantization implies that

$$[a(\vec{k}), a^\dagger(\vec{k}')] = (2\pi)^3(2\omega)\delta^3(\vec{k} - \vec{k}'),$$

creation and annihilation operators. The quantum field is a superposition of creation and annihilation operators. Note also that

$$H = \frac{1}{2} \int \frac{d^3 k}{(2\pi)^2 (2\omega)} \omega (a(\vec{k}) a^\dagger(\vec{k}) + a^\dagger(\vec{k}) a(\vec{k})).$$

Need to normal order the first term. Define $:AB:$ for operators A and B to mean that the terms are arranged so that the annihilation operators are on the right, so annihilates the vacuum.

- Causality? Compute $[\phi(x), \phi(y)] = D(x-y) - D(y-x)$, where

$$\langle 0 | \phi(x) \phi(y) | 0 \rangle = D(x-y) \equiv \int \frac{d^3 k}{(2\pi)^3 2\omega(k)} e^{-ik(x-y)}.$$

Note that the commutator is a c-number, not an operator. For spacelike separation, $(x-y)^2 = -r^2$, get $D(x-y) = \frac{m}{2\pi^2 r} K_1(mr)$. Although $D(x-y) \sim \exp(-m|\vec{x} - \vec{y}|)$ is non-vanishing outside the forward light cone, the above difference is not. Good.

- Get more interesting theories by adding interactions, e.g. $V(\phi) = \frac{1}{2} m^2 \phi^2 + \lambda \phi^4$, treat 2nd term as a perturbation.

- Symmetries of \mathcal{L} and Noether's theorem. If a variation $\delta\phi_a$ changes $\delta L = \partial_\mu F^\mu$, then it's a symmetry of the action and there is a conserved current: $j^\mu = \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi_a)} \delta\phi_a - F^\mu$.

Example: $x^\mu \rightarrow x^\mu + \epsilon^\mu$, $\delta\phi_a = \epsilon^\nu \partial_\nu \phi_a$, $\delta\mathcal{L} = \epsilon^\nu \partial_\nu \mathcal{L}$ (assuming no explicit x dependence). Get $T_{\mu\nu} = \frac{\partial \mathcal{L}}{\partial \partial_\mu \phi_a} \partial_\nu \phi_a - g_{\mu\nu} \mathcal{L}$. Stress energy tensor. Conserved charge is $P_\mu = \int d^3 \vec{x} T_{\mu 0}$.

Another example: $\Lambda_\nu^\mu = \delta_\nu^\mu + \omega_\nu^\mu$, leads to conserved $M_{\mu\rho\sigma} = x_\mu T_{\rho\sigma} - x_\sigma T_{\rho\mu}$. Conserved charge is $M_{\rho\sigma} = \int d^3 x M_{0\rho\sigma}$. Conserved angular momentum.

- Consider green's functions for the KG equation,

$$(\partial_x^2 + m^2)D(x-y) = -i\delta^4(x-y),$$

which we can use to solve $(\partial_x^2 + m^2)\phi(x) = \rho(x)$, via $\phi(x) = \phi_0(x) + i \int d^4 y D(x-y)\rho(y)$, where ϕ_0 is a solution of the homogeneous KG equation. Get

$$D(x-y) = \int \frac{d^4 k}{(2\pi)^4} \frac{i}{k^2 - m^2} e^{-ik(x-y)}.$$

Consider the k_0 integral in the complex plane. There are poles at $k_0 = \pm\omega_k$, where $\omega_k \equiv +\sqrt{\vec{k}^2 + m^2}$. There are choices about whether the contour goes above or below

the poles. Going above both poles gives the retarded green's function, $D_R(x - y)$ which vanishes for $x_0 < y_0$. For $x_0 > y_0$, note that

$$D_R = \int \frac{d^3k}{(2\pi)^3 2\omega_k} (e^{-ik(x-y)} - e^{ik(x-y)}) \equiv D(x-y) - D(y-x) = \langle [\phi(x), \phi(y)] \rangle.$$

Going below both poles gives the advanced propagator, which vanishes for $y_0 < x_0$.

- Feynman propagator. Define

$$D_F(x-y) \equiv \langle T\phi(x)\phi(y) \rangle = \begin{cases} \langle \phi(x)\phi(y) \rangle & \text{if } x_0 > y_0 \\ \langle \phi(y)\phi(x) \rangle & \text{if } y_0 > x_0 \end{cases}.$$

Here T means to time order: order operators so that earliest is on the right, to latest on left. Object like $\langle T\phi(x_1)\dots\phi(x_n) \rangle$ will play a central role in this class. Time ordering convention can be understood by considering time evolution in $\langle t_f | t_i \rangle$. Evaluate $D_F(x-y)$ by going to momentum space:

$$D_F(x-y) = \int \frac{d^4k}{(2\pi)^4} \frac{i}{k^2 - m^2 + i\epsilon} e^{-ik(x-y)},$$

where $\epsilon \rightarrow 0^+$ enforces that we go below the $-\omega_k$ pole and above the $+\omega_k$ pole, i.e. we get $D(x-y)$ if $x_0 > y_0$, and $D(y-x)$ if $x_0 < y_0$, as desired from the definition of time ordering. This ensures causality.

- Define contraction of two fields $A(x)$ and $B(y)$ by $T(A(x)B(y)) - :A(x)B(y):$. This is a number, not an operator, e.g. for $x^0 > y^0$ the contraction is $[A^+, B^-]$. So can put between vacuum states to get that the contraction is $\langle TA(x)B(y) \rangle$.

- Simple example of interacting theory:

$$\mathcal{L} = \frac{1}{2}(\partial\phi^2 - \mu^2\phi^2) + (\partial\psi^\dagger\partial\psi - m^2\psi^\dagger\psi) - g\phi\psi\psi^\dagger.$$

Toy model for interacting nucleons and mesons. Treat last term as a perturbation.

- Dyson's formula. Compute scattering S-matrices. Consider asymptotic in and out states, with the interaction turned off. Time evolve, with the interaction smoothly turned on and off in the middle.

$$|\psi(t)\rangle = T e^{-i \int d^4x \mathcal{H}_I} |i\rangle.$$

Now use Wick's theorem:

$$T(\phi_1 \dots \phi_n) = : \phi_1 \dots \phi_n : + \sum_{\text{contractions}} : \phi_1 \dots \phi_n :$$

to get rid of the time ordered products. Thereby compute probability amplitude for a given process

$$\langle f | (S - 1) | i \rangle = i \mathcal{A}_{fi} (2\pi)^4 \delta^{(4)}(p_f - p_i).$$

Next time, look at some examples, and connect with Feynman diagrams.